



proANUBIS experiment: Status and prospects for Run 3

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On behalf of ANUBIS

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Cavendish Laboratory, University of Cambridge*

Large Hadron Collider Physics conference (LHCP – 2023), University of Belgrade, Belgrade, May 22–26, 2023

Introduction

MONDAY, 22 MAY

Already....

14:30

Overview of searches, prospects, what are we missing?

Speakers: Victor Martin Lozano (IFIC/UV), Victor Martin Lozano (DESY)

14:48

LLP overview: theory perspective ¶

Speaker: Giovanna Cottin (Universidad Adolfo Ibáñez (CL))

15:42

LLP results from CMS

Speaker: Ang Li (University of Virginia (US))

16:00

LLP results from ATLAS ¶

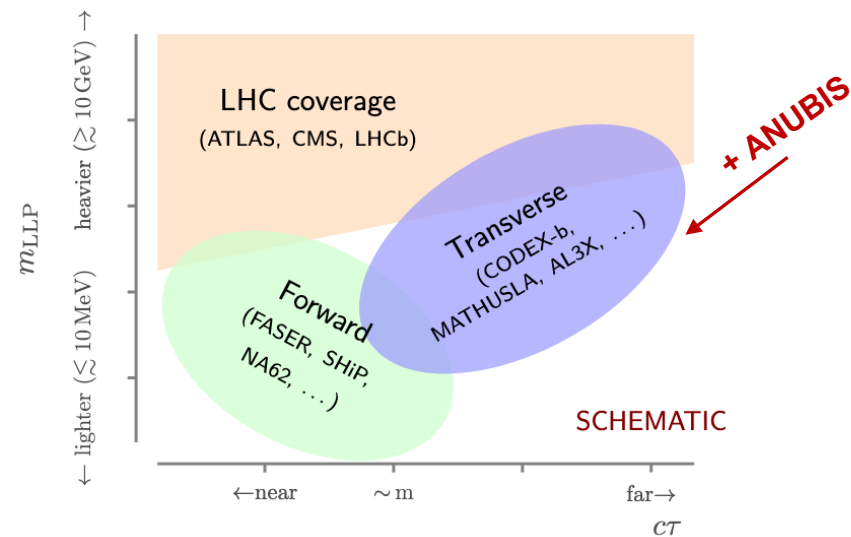
Speaker: Mohsen Naseri (Carleton University (CA))

Where could New Physics (NP) be hiding?

- A logical possibility is that conventional collider detectors are not optimized to detect the NP if it exists (e.g., LLP's)
- **Design experiments carefully?**
 - > LLPs characterized by their **long lifetimes**, and might be decaying at distances **beyond the detector's reach**, that poses a challenge as traditional collider detectors are optimized for detecting prompt, short-lived particles
 - > **Comprehensive detector coverage (or dedicated detectors)** could be effective in identifying and characterizing LLP events

Dedicated detectors

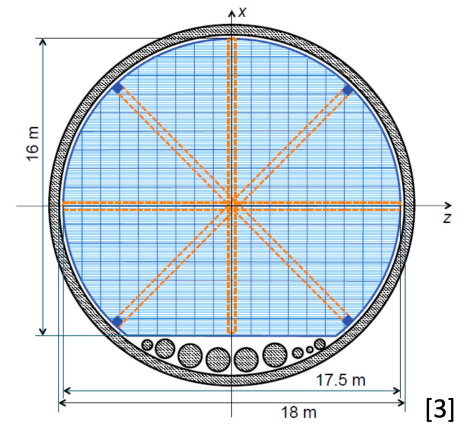
- Low-energy and fixed-target experiments like SHiP, NA62, SEAQUEST, etc. might help but many predicted LLPs can only be produced in collision energies such as at LHC
- The LHC could produce many LLPs with MeV - TeV masses that cannot be produced anywhere else, but that existing detectors cannot discover
- Augmenting its capabilities with relatively modestly-priced external detectors to maximize the discovery potential for new physics should be a high-priority goal
- Dedicated LLP Detectors
 - > Several new proposals to address the significant gap in the LHC's reach for long-lived particles
 - > Motivations for construction of **dedicated detectors** further away from the interaction points.
 - > These are of two main types
 - On-axis:** these have increased sensitivity to lighter LLPs (FASER, MAPP)
 - Off-axis,** these have increased sensitivity to heavier LLPs (MATHUSLA, CODEX-b, ALX3, ANUBIS)
- Use existing LHC infrastructure (and detector technology) along the beamline to cut down the major civil engineering (and R&D) costs



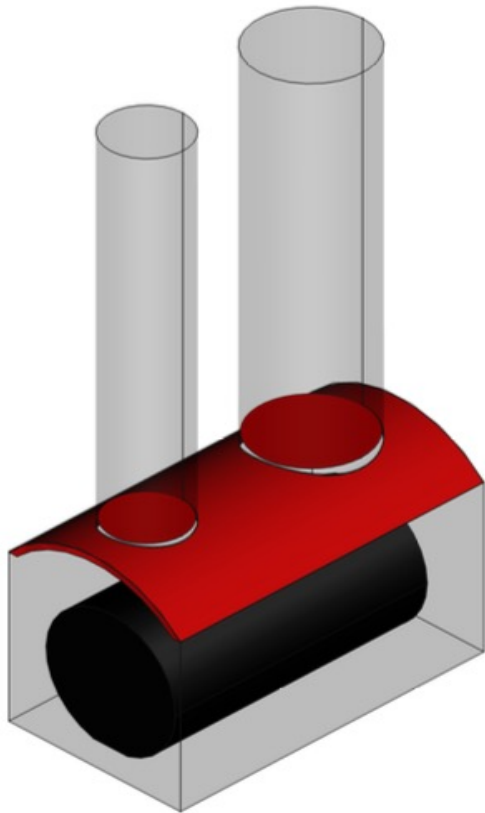
Dedicated detectors: ANUBIS

ANUBIS – AN Underground Belayed In-Shaft search experiment

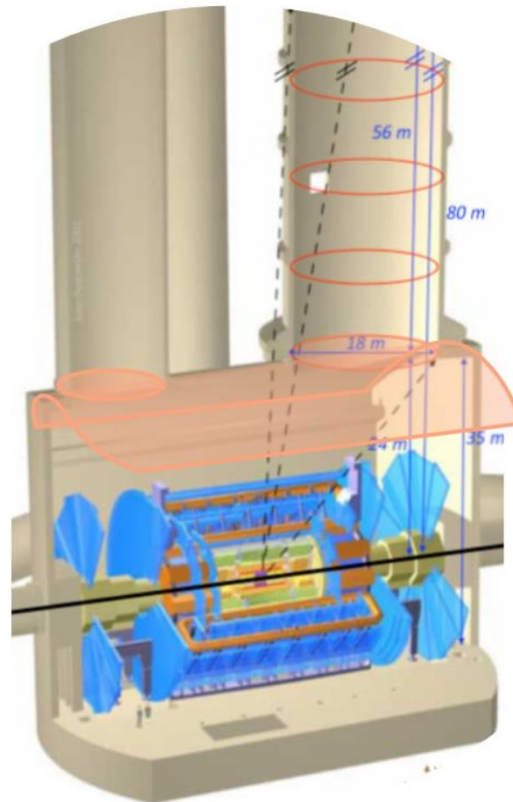
- Proposal to instrument the ceiling of the ATLAS Cavern at Point-1
 - > Include stations in the two service shafts (PX14, PX16)
 - > Ceiling approximately 20m away from the ATLAS IP
 - > Cavern ceiling proposal shown to be more sensitive (compared to Shaft only)
 - > Larger active volume ($4.3 \times 10^4 \text{ m}^3$ vs $1.3 \times 10^4 \text{ m}^3$) and large detector area $\sim 10^3 \text{ m}^2$



PX14 Shaft: Cross sectional view



Proposal



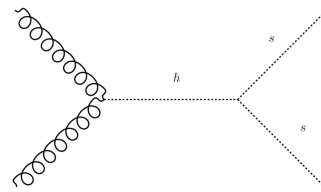
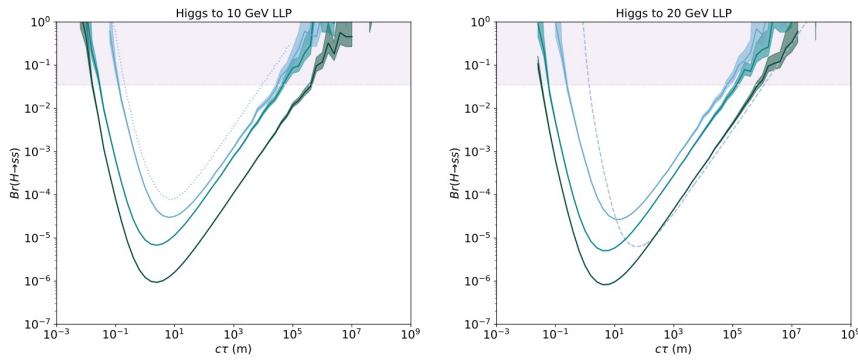
arXiv:1909.13022

<https://twiki.cern.ch/twiki/bin/view/ANUBIS>



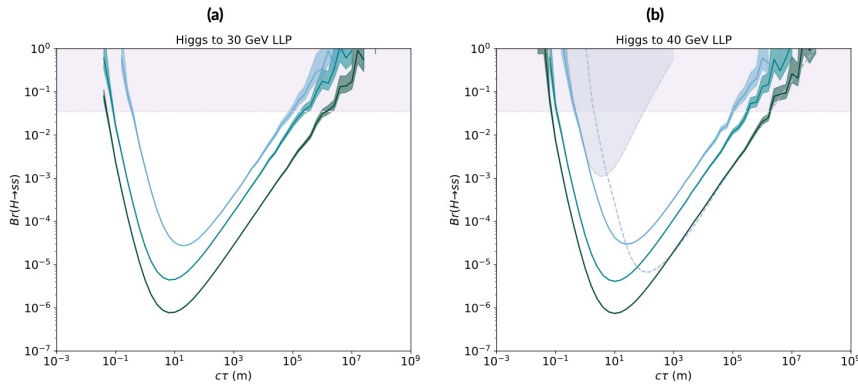
PX14 Shaft + Ceiling

Sensitivity and expected backgrounds



Benchmark signal

- Projected Sensitivity
- Estimates from simple (non-G4) simulation
- Sensitivity assumed from 90 observed events for the ceiling geometry



(c) @Toby Satterthwaite (d)

- ANUBIS ceiling
- ANUBIS PX14 shaft -- cavern or shaft decay
- ANUBIS PX14 shaft -- shaft decay
- ANUBIS sensitivity $\pm 1\sigma$
- ⋯ CODEX-b ($\mathcal{L} = 1 \text{ ab}^{-1}$)
- - - MATHUSLA ($\sqrt{s} = 14 \text{ TeV}$, $\mathcal{L} = 3 \text{ ab}^{-1}$)
- ATLAS limit ($\sqrt{s} = 13 \text{ TeV}$, $\mathcal{L} = 36.1 \text{ fb}^{-1}$)
- $H \rightarrow$ Invisible limit ($\sqrt{s} = 13 \text{ TeV}$, $\mathcal{L} = 3 \text{ ab}^{-1}$)

Backgrounds

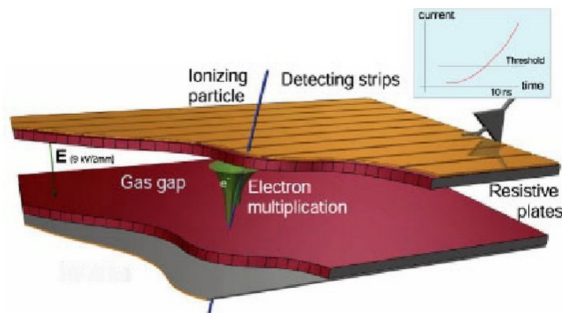
- Main background comes from neutron-air interactions, K_L^0 decays and interactions
- ATLAS Calorimeter acts as an active veto
10 interactions lengths $\rightarrow 10^{-5}$ reduction in rate
- Further reduction from ATLAS-level selections
 $E_{T, \text{miss}}$
Isolation \rightarrow hadronic particles produced as part of jets
- Exploit good RPC timing resolution to distinguish the SM backgrounds and heavier LLP's
- SM backgrounds are light and would be travelling $\sim c$ to pass the specific selections, while as heavier LLP's should arrive later

Detailed Simulations are being performed @Jon Burr

ANUBIS - detector technology

Detector requirements/technology

- Very large detector area required to cover
- **Good time resolution**
- Angular resolution limits ability to measure collimated decays
 $\delta\alpha < 0.01\text{rad} \Rightarrow m_{\text{LLP}} > 0.5 \text{ GeV}$
 Improves ability to localize the vertex
- **High hit efficiency to avoid missing events and it also improves discrimination between signal and background**
- Motivates use of RPCs
 - > Known technology and very simple and offers cost effective solution



Traditional RPC design

- Next generation of RPC → BIS78 technology
 - Higher rate capability → kHz/cm²
 - Longer longevity: >10 years @ HL-LHC
 - Higher spatial resolution: <1 cm
 - Higher time resolution: ~ 0.5 ns

I. Required performance specifications for ANUBIS.

Parameter	Specification
Time resolution	$\delta t \lesssim 0.5 \text{ ns}$
Angular resolution	$\delta\alpha \lesssim 0.01 \text{ rad}$
Spatial resolution	$\delta x, \delta z \lesssim 0.5 \text{ cm}$
Per-layer hit efficiency	$\varepsilon \gtrsim 98\%$

Parameters (detector + FE boards)

	Standard RPC	BIS78 RPC
FEE		
Effective threshold	1mV	0.5mV
Power consumption	30 mW	6 mW
Technology	GaAs	BJT Si + SiGe
Discriminator	Embedded	Separated
TDC embedded	No	No
Detector		
Gap Width	2 mm	1 mm
Operating voltage	9600 V	5800 V
Electrode thickness	1.8 mm	1.2 mm
Time resolution	1 ns	0.4 ns

proANUBIS - prototype of ANUBIS

Proof-of-Concept - demonstrator detector for ANUBIS

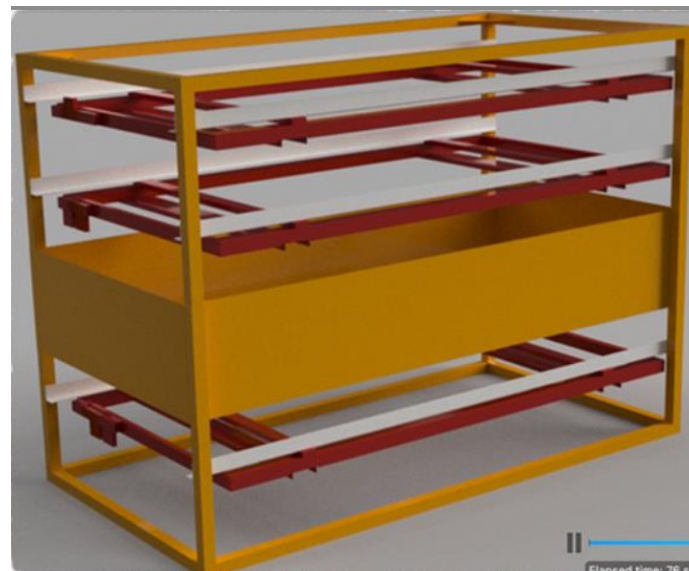
- Form three tracking station using BIS78 RPC chambers and install it in the cavern during 2022/23
- **Detector performance and Physics goals**

Hit/track efficiency

Identify muons selected by ATLAS triggers and synchronize the detectors

Validate Geant4 Simulations

Measure rates of hadrons from punch through jets



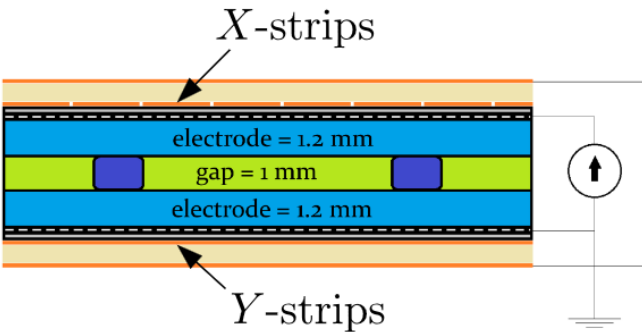
Design of demonstrator/prototype detector (to be) installed during YETS23

proANUBIS - detector construction

- Full process of construction, started last summer from the panels, soldering terminators, testing FE boards and attaching them to panels
- Uses new generation of BIS78 triplet RPCs from ATLAS muon phase-I Upgrade



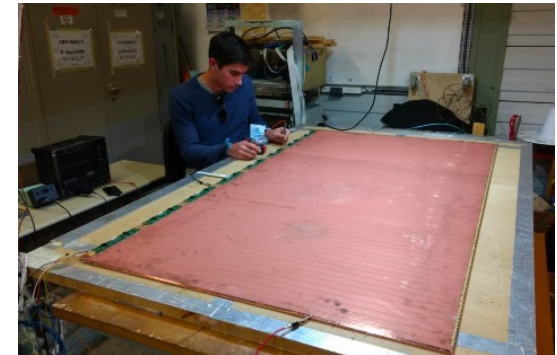
FE boards testing @Jon, Toby, and Oleg



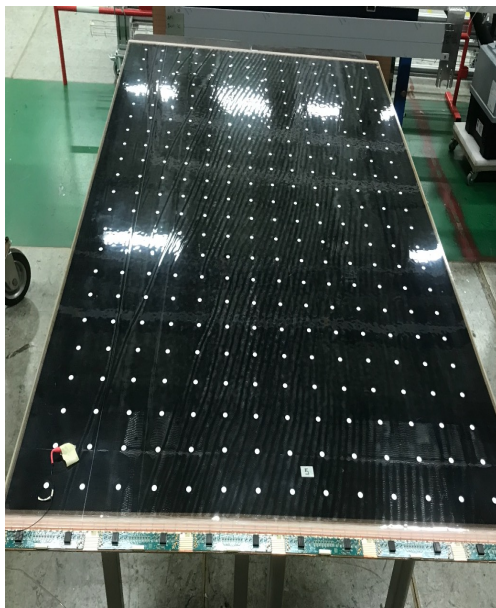
[RPC singlet] RPC Singlet design



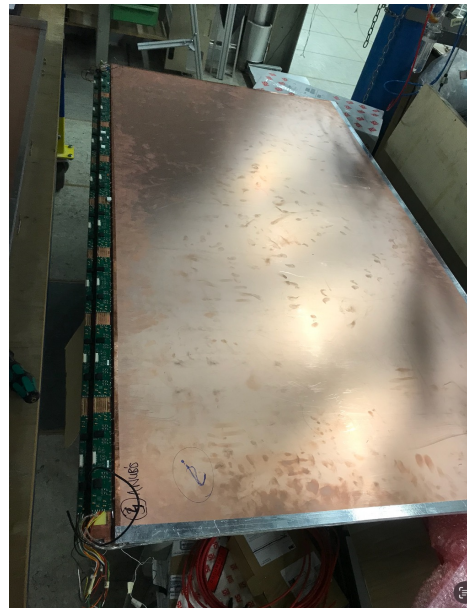
PCB strip panel for BIS78



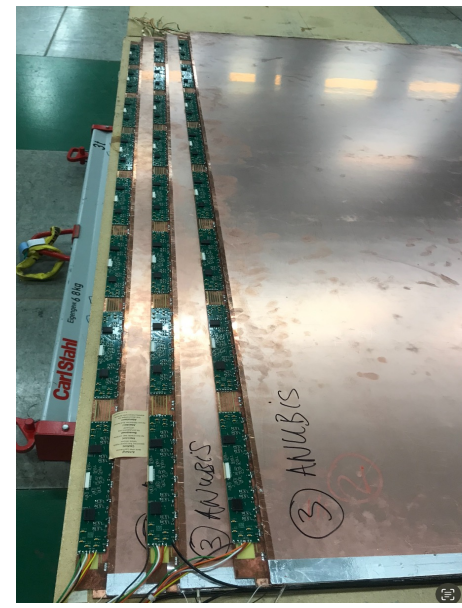
Strip panel testing @Toby



RPC gas gap

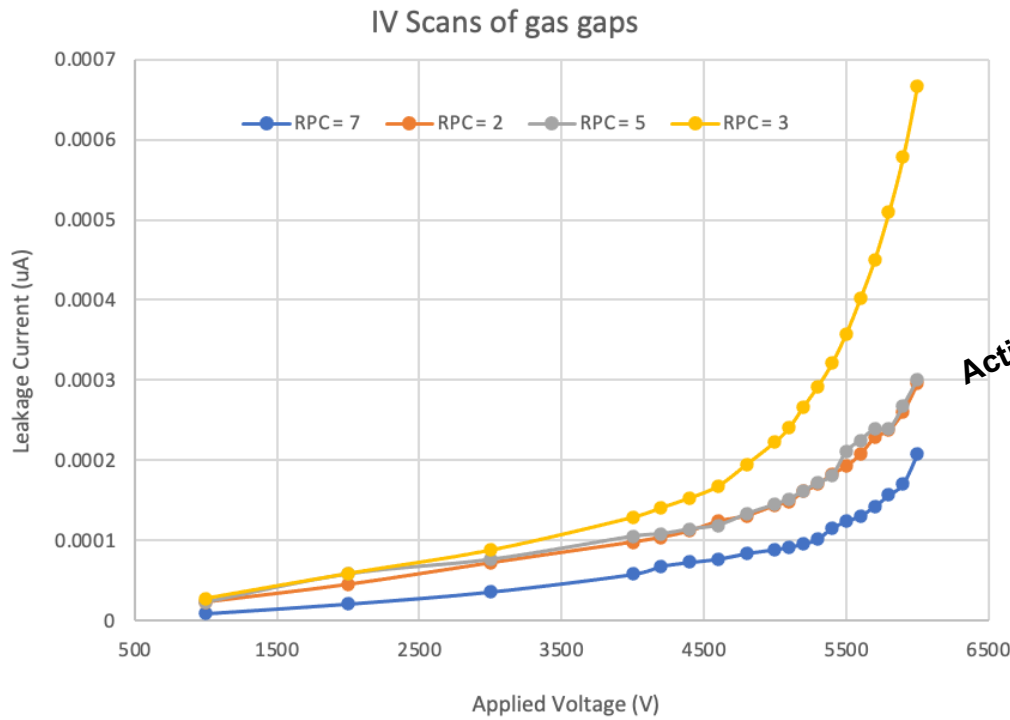


RPC Singlet (assembled one)

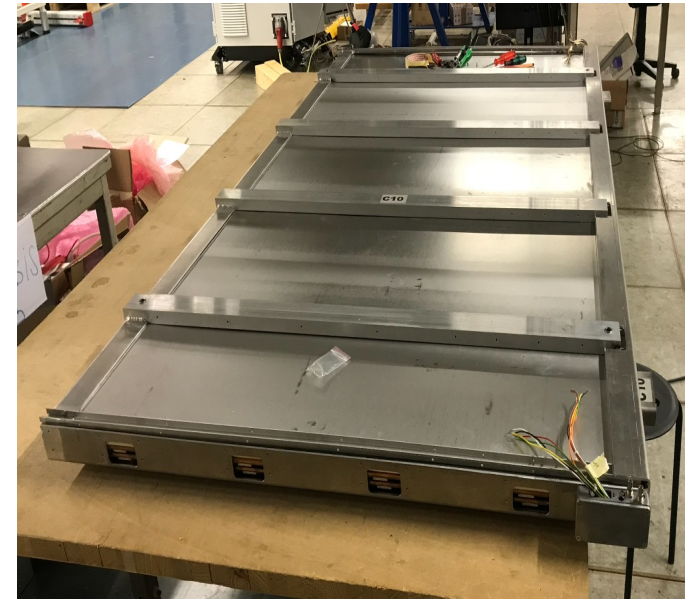


3 RPCs waiting for integration

proANUBIS - detector construction



Activities @BB5



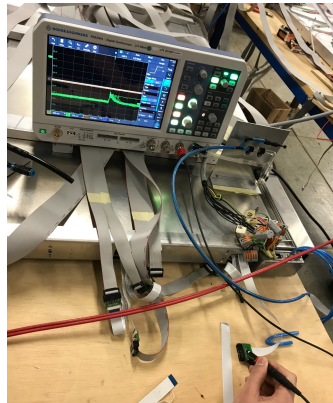
Fully integrated RPC doublet

- Everything wasn't so smooth: **debugging and fixing**

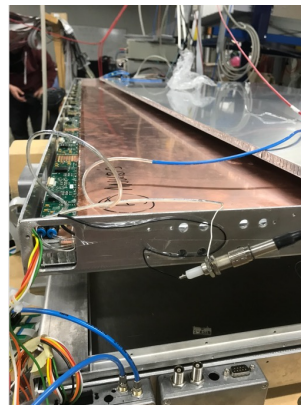
- For proANUBIS
Triplet, Singlet, Doublet



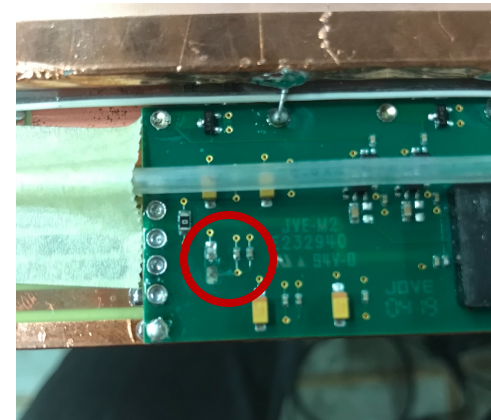
Triplet and doublet integrated by now but none worked perfectly



Debugging dead channels



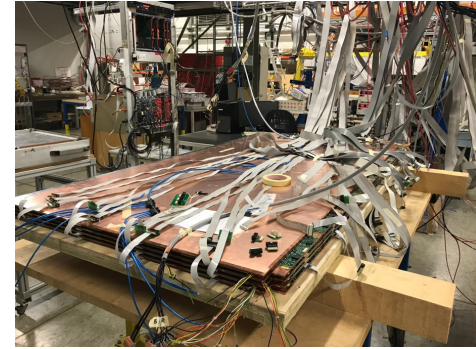
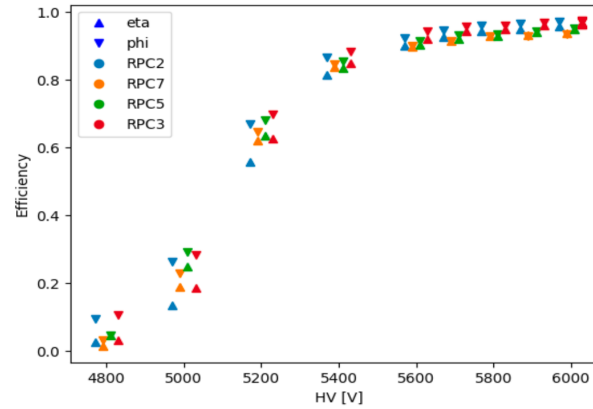
Debugging HV short



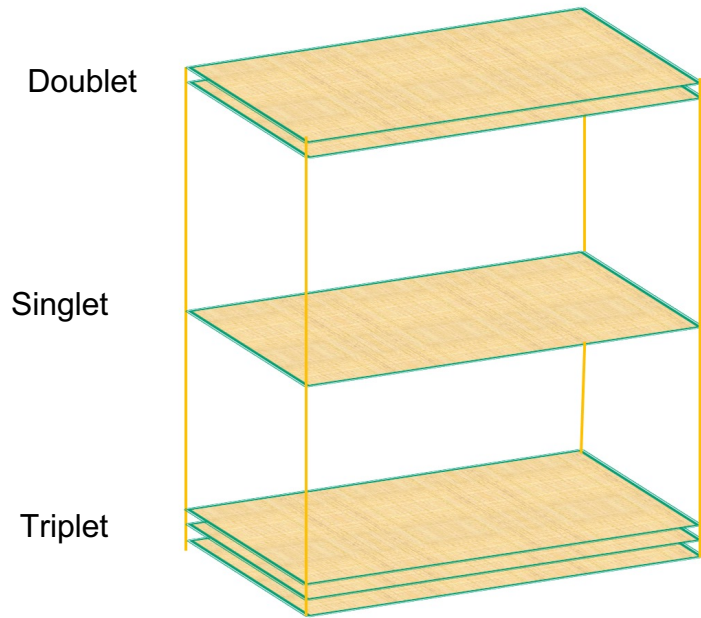
Replacing FE board/s to fix dead channels

proANUBIS detector - Efficiency measurements

- Efficiency has been estimated using cosmic muons
- All RPC's behaving very well with good performance
- Three integrated Chambers
 - Singlet – consists of one RPC detector
 - Doublet – consists of two RPC detectors
 - Triplet – consists of three RPC detectors



Efficiency measurements of individual detectors before integration @BB5



Design of proANUBIS detector



Preparing setup (RPC singlet + doublet) for efficiency measurement after integration @BB5



Set up for efficiency measurements after integration @BB5

proANUBIS - exciting scenes during installation - from BB5 to experimental cavern (YETS-2022)



Frame lifted by Crane @BB5



Chamber integration @BB5



Fully integrated system being tested @BB5



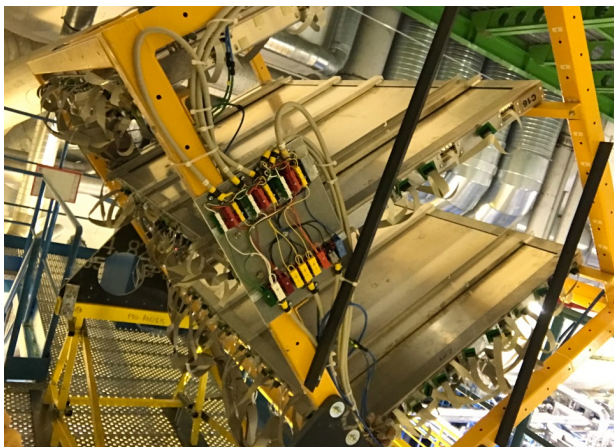
proANUBIS being lifted by Crane to pass it over through PX14 @SX15



proANUBIS lowering in the ATLAS cavern



proANUBIS base stand



proANUBIS closeup view after installation

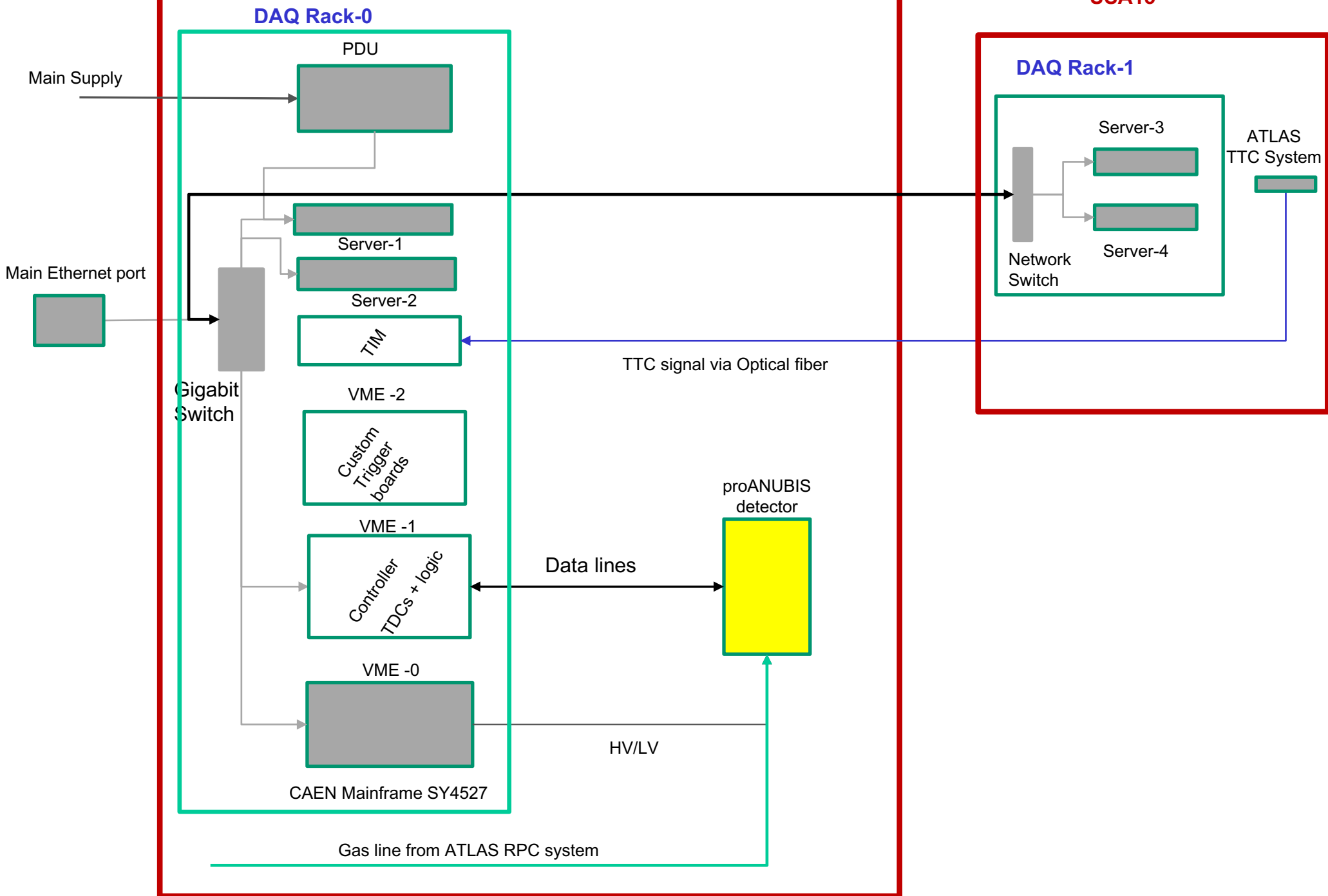


proANUBIS + DAQ rack installed in their positions within the ATLAS experimental Cavern (Level 12 of UX15)

proANUBIS setup

Experimental cavern

USA15



proANUBIS - Close-up look at DAQ

DAQ Rack-0



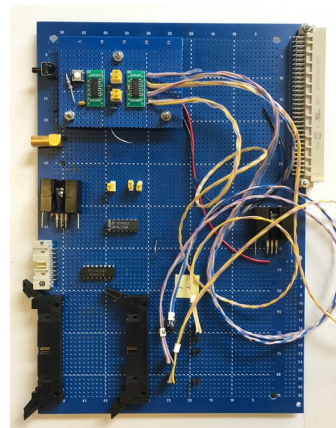
DAQ rack: without cabling



VME-1



VME-0



Custom board used for NIM/ECL conversion @Rick Shaw

VME-0 (6U)

- Controller CAEN V4718
- TDC CAEN V767
- Custom NIM to ECL Con. board
- Custom Trigger board
- Custom Majority OR board

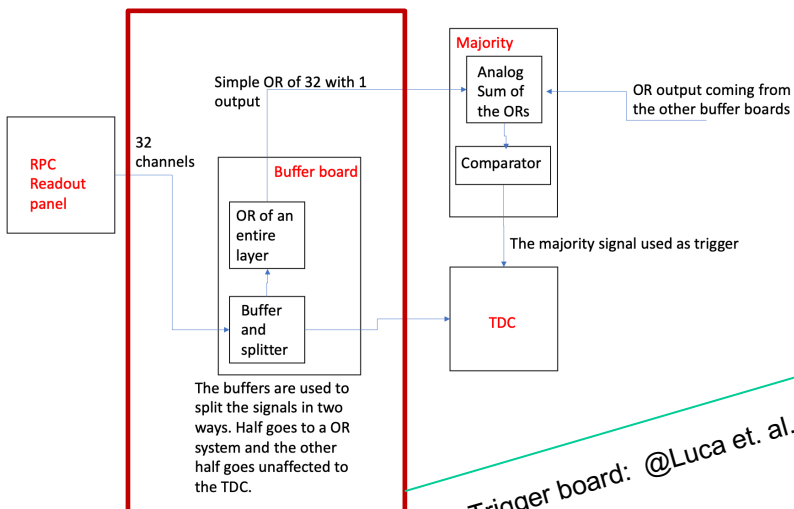
VME-1 (6U)

- Custom Trigger boards

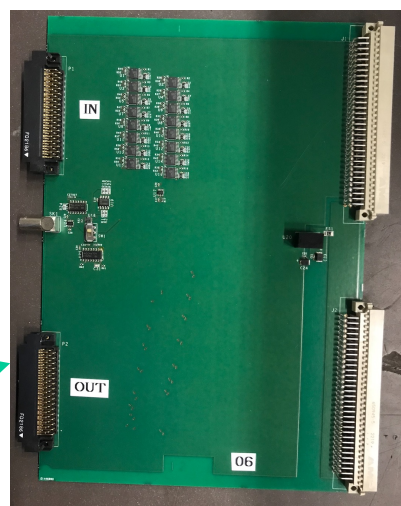
VME-2 (9U)

- TIM card for TTC (Next slide)

Trigger Logic



Custom Trigger board: @Luca et al.,



@Rick et al.,

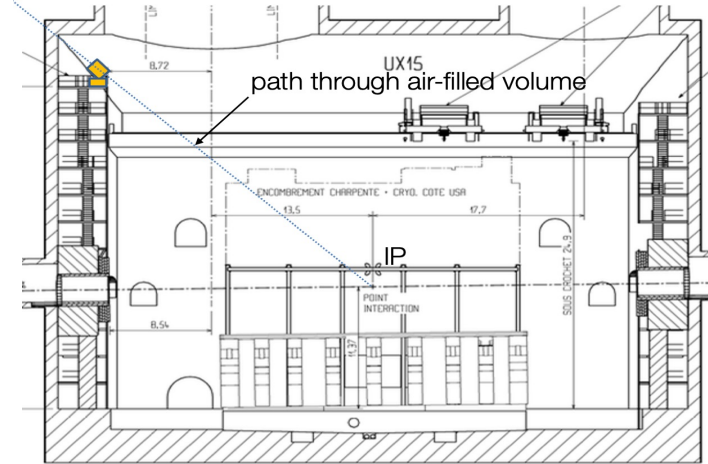


DAQ rack: fully cabled-up

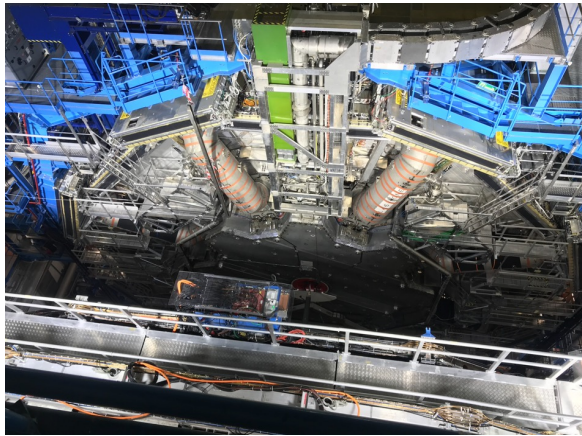
proANUBIS - detector location



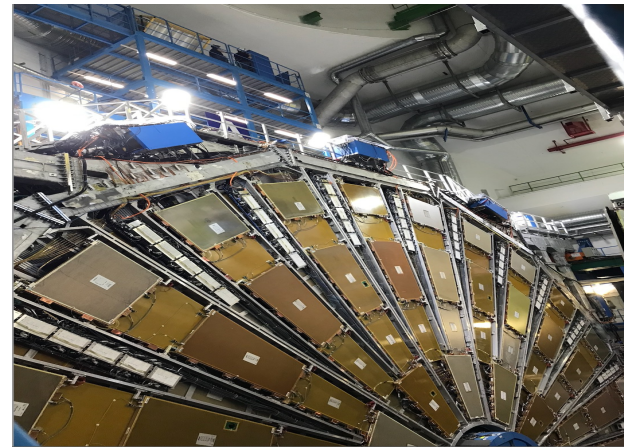
proANUBIS (DAQ + Detector) after installation



proANUBIS location @UX15



Downward-view (towards IP) from detector @UX15

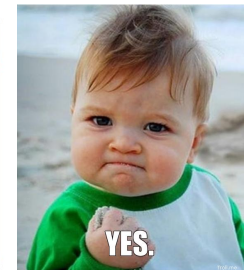
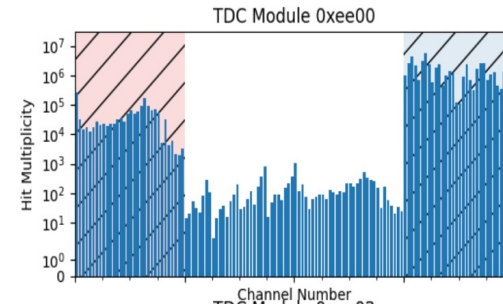


Upward view (close to beam pipe) @UX15

First glimpse at cosmic data!

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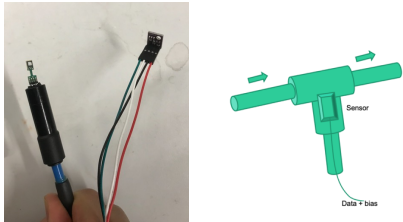
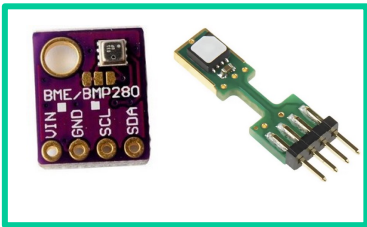
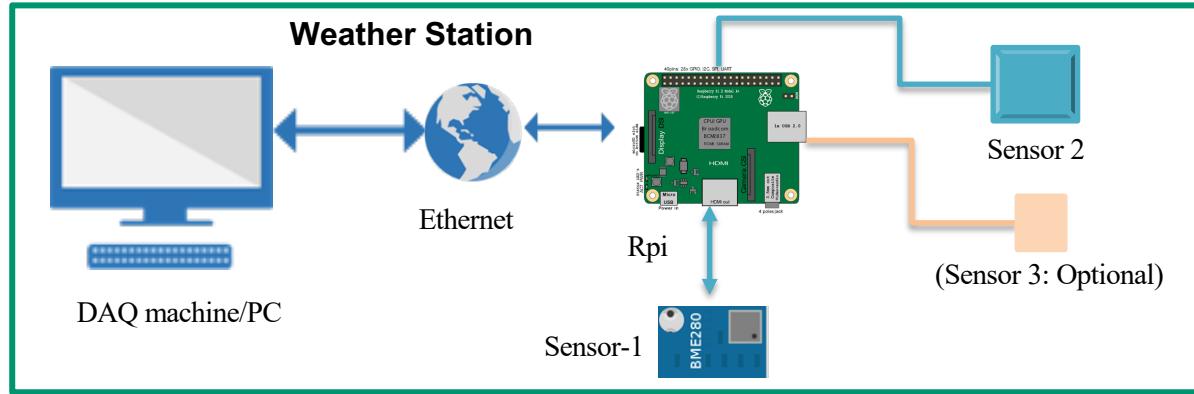
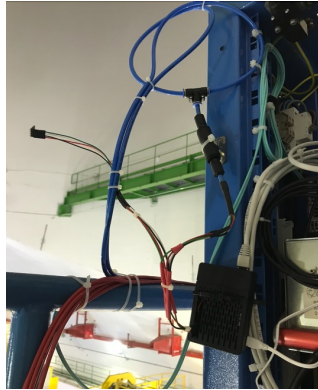
Reading Buffer: F84001A6; Byte Stream: 1111100001000000000000110100110; Header with Event No.: 422;
Reading Buffer: 4F100614; Byte Stream: 01001111000100000000011000010100; Channel No.: 79; Time Measurement: 1244.80 ns;
Reading Buffer: F9200001; Byte Stream: 11111001001000000000000000000000; End of Block (EOB)
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Reading Buffer: 471005E7; Byte Stream: 0100011100010000000000010111100111; Channel No.: 71; Time Measurement: 1208.80 ns;
Reading Buffer: 491005F0; Byte Stream: 01001001000100000000010111110000; Channel No.: 73; Time Measurement: 1216.00 ns;
Reading Buffer: 4A1005EE; Byte Stream: 01001010000100000000010111101110; Channel No.: 74; Time Measurement: 1214.40 ns;
Reading Buffer: 48100626; Byte Stream: 01001000000100000000011000100110; Channel No.: 72; Time Measurement: 1259.20 ns;
Reading Buffer: F9200004; Byte Stream: 11111001001000000000000000000100; End of Block (EOB)
Reading Buffer: F84001A8; Byte Stream: 11111000010000000000000110101000; Header with Event No.: 424;
Reading Buffer: F9200000; Byte Stream: 11111001001000000000000000000000; End of Block (EOB)
Reading Buffer: F84001A9; Byte Stream: 11111000010000000000000110101001; Header with Event No.: 425;
Reading Buffer: 0010021F; Byte Stream: 000000000001000000000001000011111; Channel No.: 0; Time Measurement: 434.40 ns;
Reading Buffer: 0010024B; Byte Stream: 000000000001000000000001001001011; Channel No.: 0; Time Measurement: 469.60 ns;
Reading Buffer: 001002A4; Byte Stream: 000000000001000000000001010100100; Channel No.: 0; Time Measurement: 540.80 ns;
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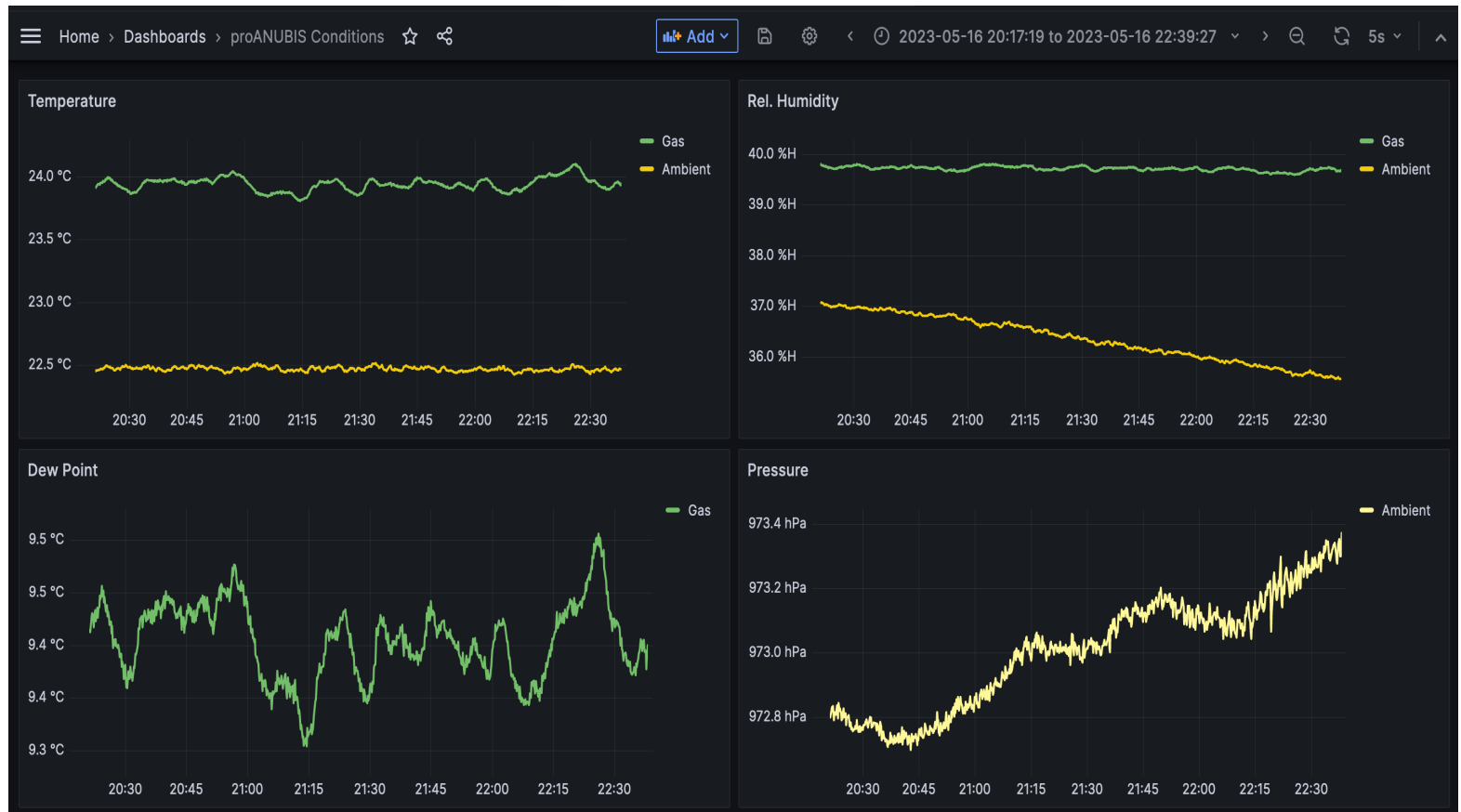
Preliminary...

Monitoring gas/ambient (T, P, Rh) conditions

- Monitoring T, P, Rh are very important for determining the performance of the proANUBIS RPC's
- Developed/installed a Weather station using commercial components



Sensors and installation



Conditions on Grafana pannel

Summary

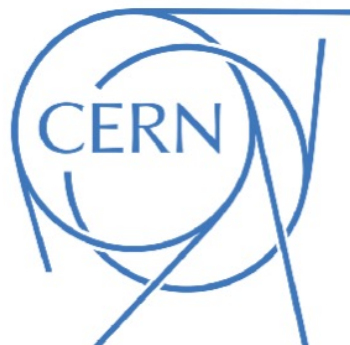
Achieved

- proANUBIS has been installed/commissioned successfully
 - > Can be fully operated remotely (LV/HV, powering cycling of servers and other individual components/modules except VME crates)
 - > Detector stable in terms of LV/HV and is running continuously
 - > Can measure ambient/gas conditions which are required for the detector performance
 - > Able to collect the cosmic data
 - > Preparing for the accumulation of the collision data ...

Current focus

- Focusing on the improvements of the DAQ - we had some (minor) issues and are following on it
- Working on to develop particle tracking software
- Near term goals are to determine
 - > Detector performance using collision data
 - Occupancy rate due to cavern background radiation
 - > Hit/track efficiency, correlate information with the ATLAS
 - > Validate background simulations
- News and recent updates: <https://twiki.cern.ch/twiki/bin/view/ANUBIS/>
- proANUBIS is running and ANUBIS is growing....!!!
- Interested to join, get in touch: anubis-active@cern.ch or oleg.brandt@cern.ch





Max-Planck-Institut für Physik
(Werner-Heisenberg-Institut)

Thank you!

Back up

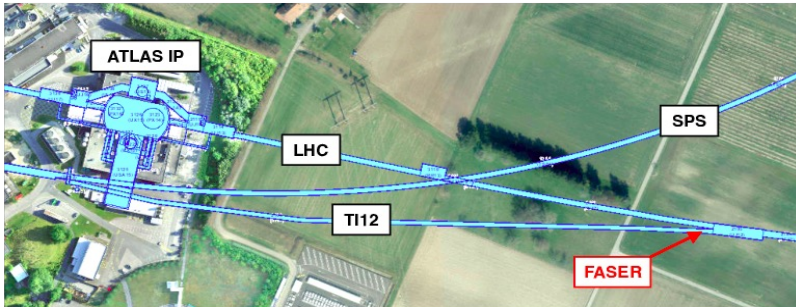
Introduction

- The discovery of the Higgs boson at the LHC and the subsequent study of its properties has greatly advanced our understanding of electroweak symmetry breaking
- Compelling evidence of **New Physics (NP)** beyond the Standard Model is still elusive
- Motivations for new NP remain strong, as many fundamental mysteries in our universe are yet to be explained by the Standard Model
- The observation of **Long-Lived Particles (LLPs)** at the LHC could provide a window into physics beyond the Standard Model
- LLPs offer a potential explanation for numerous open issues in our understanding of the universe
 - The Hierarchy Problem
 - Dark Matter
 - Neutrino Masses
 - The Baryon Asymmetry of the Universe
- The search for LLPs at the LHC involves dedicated analysis techniques to enhance their sensitivity and/or to identify their unique signatures
- Strategies for LLP searches at the LHC:
 - **Delayed Timing:** LLPs with **long lifetimes** can be identified by delayed timing between their production and decay
 - **Displaced Vertices:** LLPs can produce secondary vertices away from the primary interaction point, indicating their presence
 - **Energy Loss:** Some LLPs interact weakly with matter, resulting in observable energy loss and distinct tracks in the detector
 - **Calorimeter and Muon Systems:** Specific detector components are designed to efficiently capture LLP signatures, such as electromagnetic and hadronic calorimeters, as well as **muon detectors**

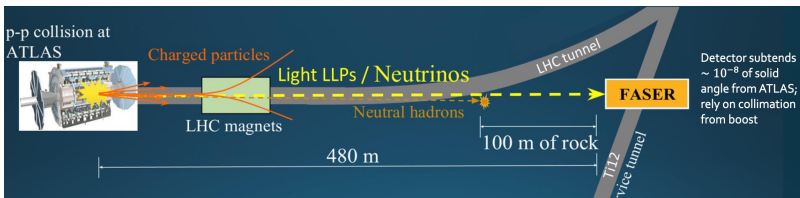
Dedicated detectors

- Two on-axis detectors constructed at the LHC: **FASER and MAPP**
- High boost of light LLPs

Signal is more focused
Much higher acceptance for a given solid angle

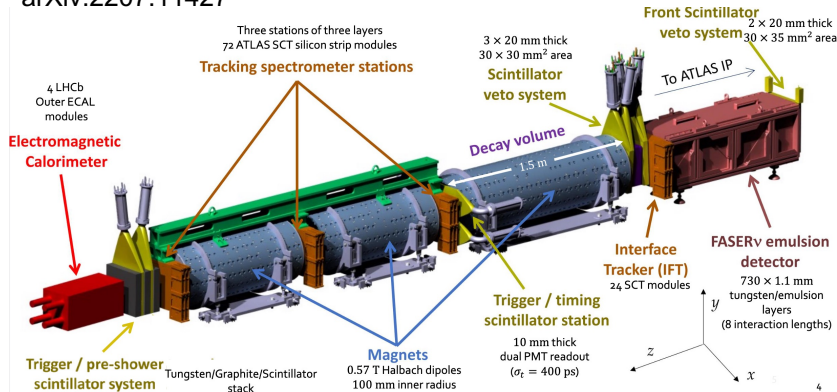


ForwArD Search ExpeRiment (FASER)

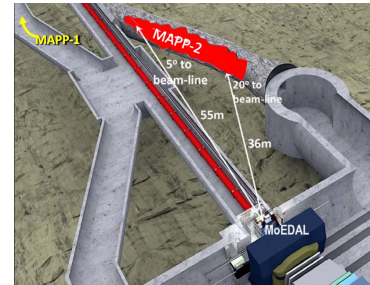


$$pp \rightarrow \text{LLP} + X, \quad \text{LLP travels } \sim 480 \text{ m}, \quad \text{LLP} \rightarrow e^+e^-, \mu^+\mu^-, \pi^+\pi^-, \gamma\gamma, \dots$$

arXiv:2207.11427

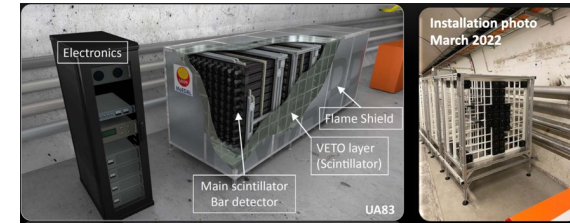


Designed to be sensitive to new-physics-induced signal events from decays of LLPs ($m \lesssim 1 \text{ GeV}$.)

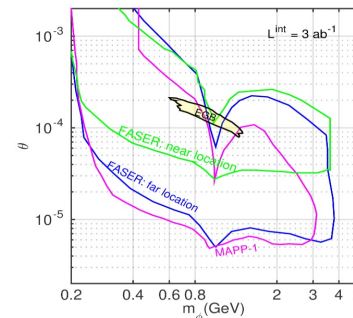


MoEDAL Apparatus for Penetrating Particles (MAPP-1)

CERN-LHCC-2021-024 / LHCC-P-022



- A new subdetector at the MoEDAL experiment in UA83, a bypass tunnel adjacent to IP8 (LHC Point 8)
- 400 scintillator bars ($10 \times 10 \times 75 \text{ cm}^3$) in 4 sections readout by PMTs
- Sensitivity to weakly interacting neutral long-lived particles



FASER and MAPP reach for dark Higgs boson for an integrated luminosity of 3 ab^{-1} at 13 TeV LHC.

arXiv:2110.09392v1

More...

12:06 **MilliQan and MoEDAL-MAPP status and prospects**

Speaker: Hualin Mei (Univ. of California Santa Barbara (US))

12:24 **FASER status and prospects**

Speaker: Noshin Tarannum (Universite de Geneve (CH))

13:00 **MoEDAL-MAPP - Detectors specialised for LLP searches**

Speaker: Dr Vasiliki Mitsou (Univ. of Valencia and CSIC (ES))

12:24 **FASER Upgrades**

Speaker: Stefano Zambito (University of Geneva)

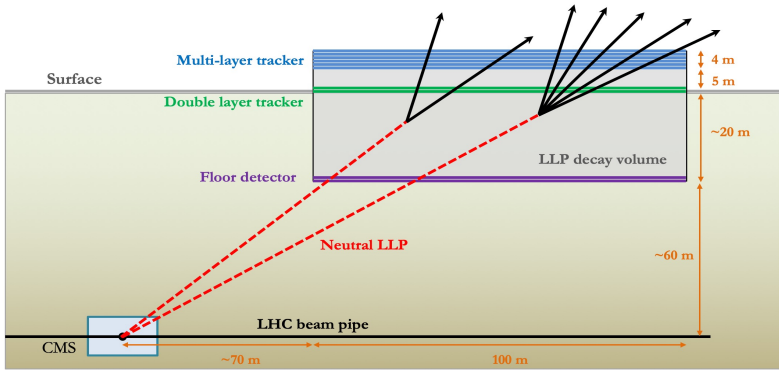
TUESDAY, 23 MAY

WEDNESDAY, 24 MAY

THURSDAY, 25 MAY

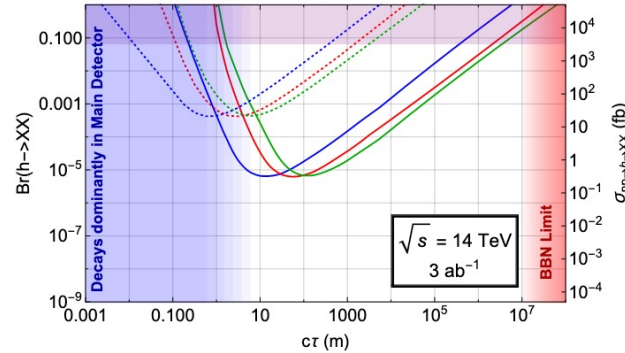
Dedicated detectors

Off-axis detectors, **MATHUSLA**, **CODEX-b**, ALX3, **ANUBIS**



MATHUSLA detector layout with a decay volume of 200 m × 200 m × 20 m

MAssive Timing Hodoscope for Ultra-Stable neutral pArticles (MATHUSLA)



MATHUSLA to address the significant gap in the LHC's reach for long-lived particles

A proposed large-scale surface detector located above CMS can detect LLPs with lifetimes near the cosmological limit of 0.1 s

12:42

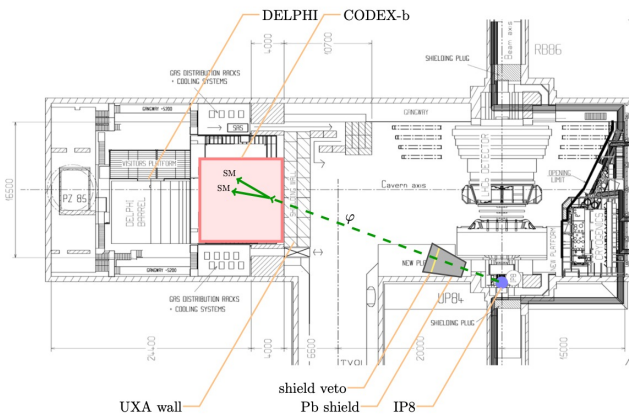
MATHUSLA

WEDNESDAY, 24 MAY

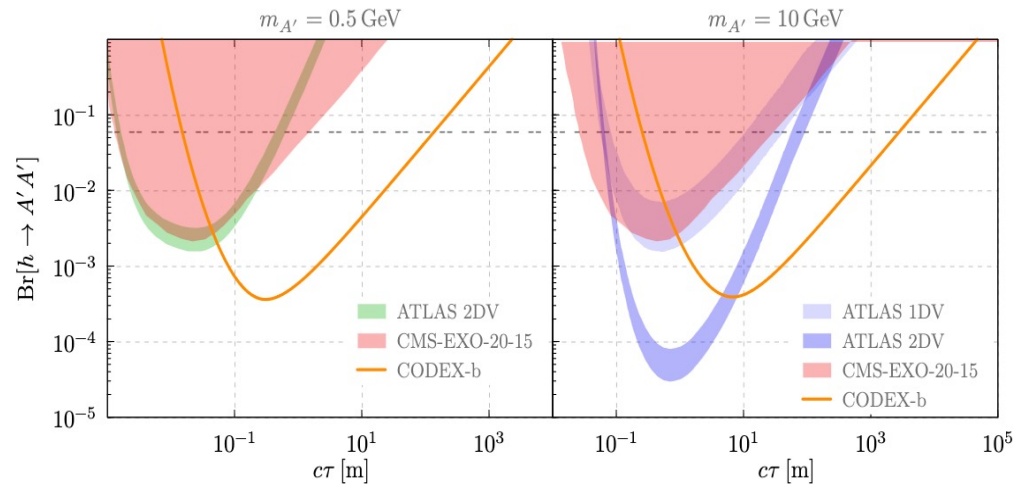
Speaker: Mason Proffitt (University of Washington (US))

COmpact Detector for EXotics at LHCb (CODEX-b)

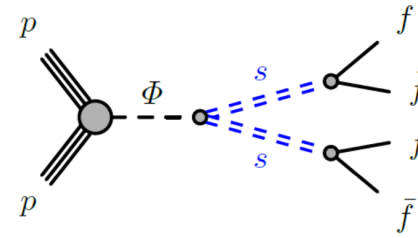
The proposed CODEX-b detector would be located roughly 25 meters from the LHCb interaction point (IP8) and have a nominal fiducial volume of 10 × 10 × 10 m³



Layout of the LHCb experimental cavern UX85 at point 8 of the LHC, overlaid with the CODEX-b volume



Example Signal



- Use to illustrate some of the main considerations
- Standard model Higgs decaying to two LLPs which then decay $s \rightarrow b\bar{b}$
- Consider a range of LLP lifetimes and masses
- s is electrically neutral and does not interact strongly
- Physics signature is a vertex which appears between two tracking layers
- SM backgrounds: neutral, long-lived
 - Neutron – air interactions
 - K_L^0 - decays and air interactions



3

Backgrounds

- Main background comes from neutron-air interactions and K_L^0 decays and interactions
- ATLAS calorimeter acts as an active veto
 - 10 hadronic interaction lengths $\Rightarrow \sim 10^{-5}$ reduction in rate
- Further reduction from ATLAS-level selections
 - E_T^{miss}
 - Isolation \Rightarrow hadronic particles usually produced as part of jets

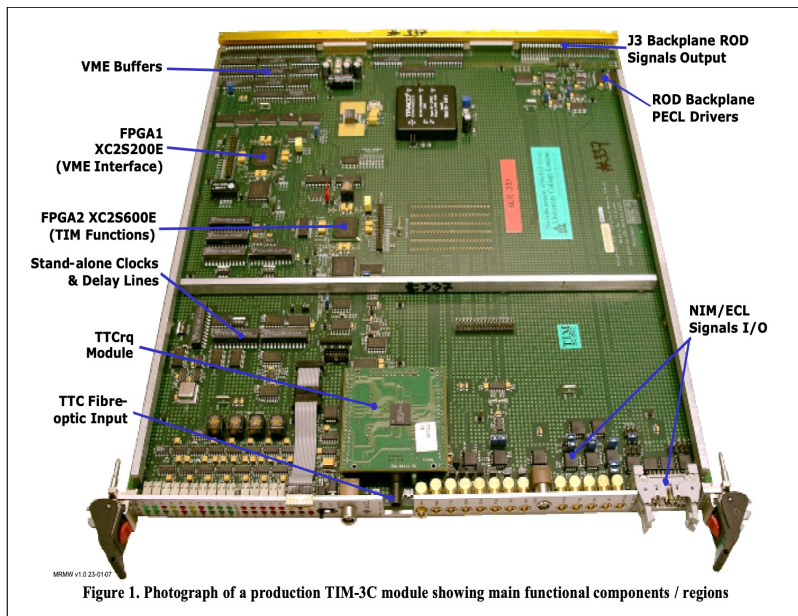
Courtesy: Jon Burr

proANUBIS - Integration with the LHC/ATLAS

- proANUBIS Detector receives TTC from LHC/ATLAS through TIM (9U-sized custom module in a VME64x crate)
- The TIM is the standard SCT/PIXEL detector interface module to the ATLAS Level-1 Trigger through the LHC- standard TTC system
- TIM (TTC Interface Module) transmits/receives the clock, fast commands, trigger, and event ID from the TTC system (TTC is the LHC-standard for Timing, Trigger and Control system)
- The optical TTC signals are received by a receiver section containing a standard TTCrx receiver chip, which decodes the TTC information into electrical form
- It does source and drive all the timing and trigger information

The TTC information, required by the RODs and by the SCT or PIXEL FE (Front End) electronics, is the following :

Clock :	BC	Bunch Crossing clock
Fast command :	L1A	Level-1 Accept
	ECR	Event Counter Reset
	BCR	Bunch Counter Reset
	CAL	Calibrate signal
Event ID :	L1ID	24-bit Level-1 trigger number
	BCID	12-bit Bunch Crossing number
	TTID	8-bit Trigger Type (+ 2 spare bits)

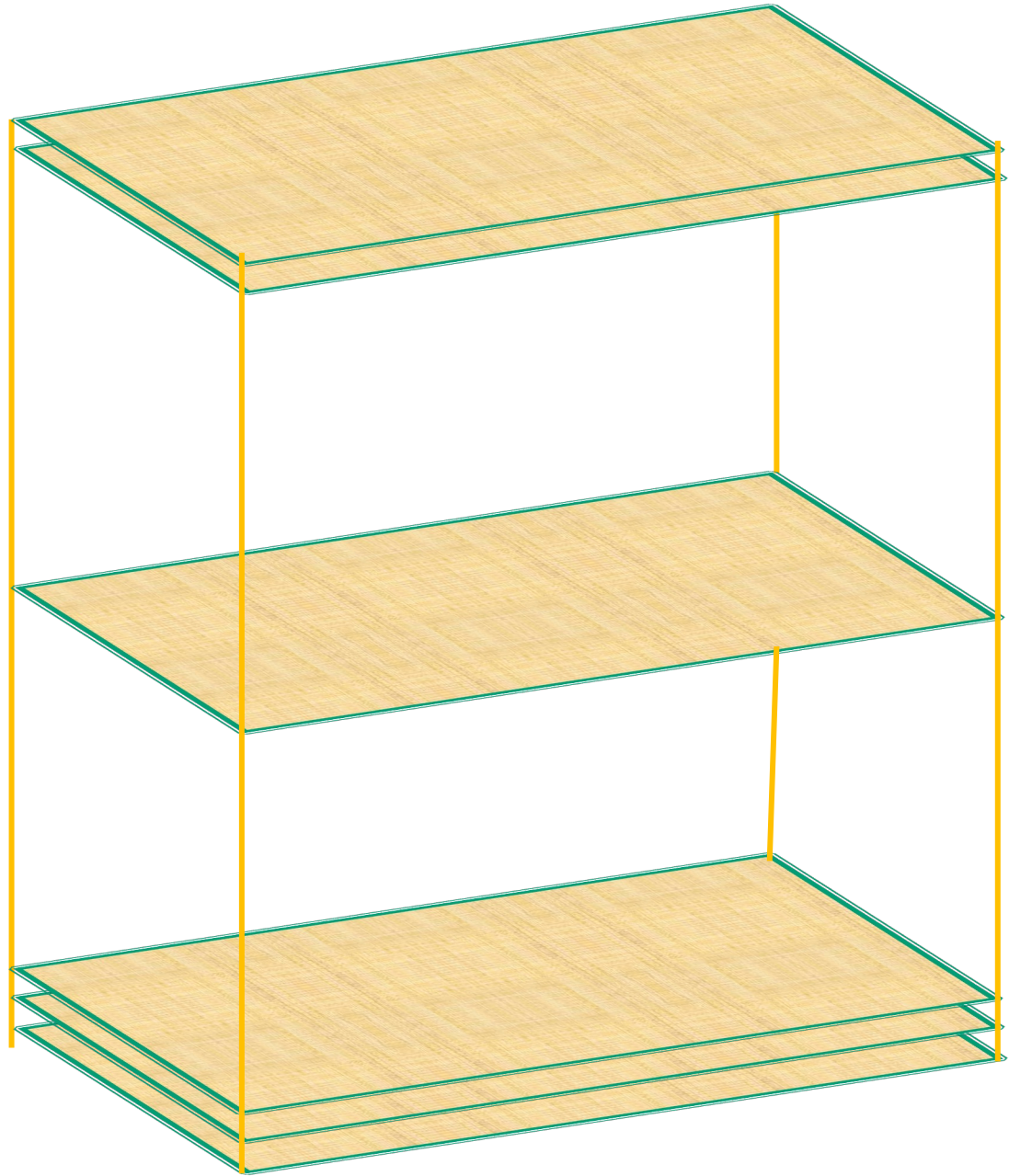
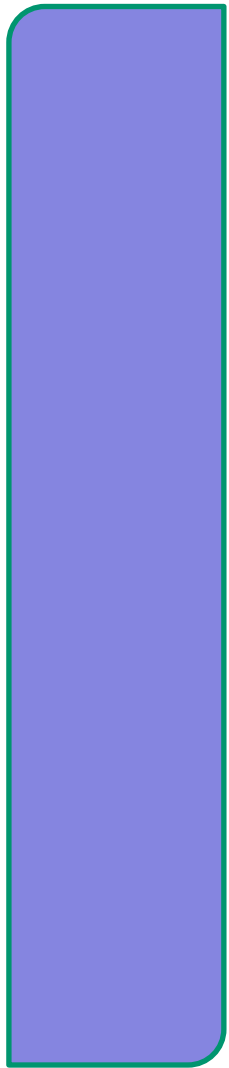


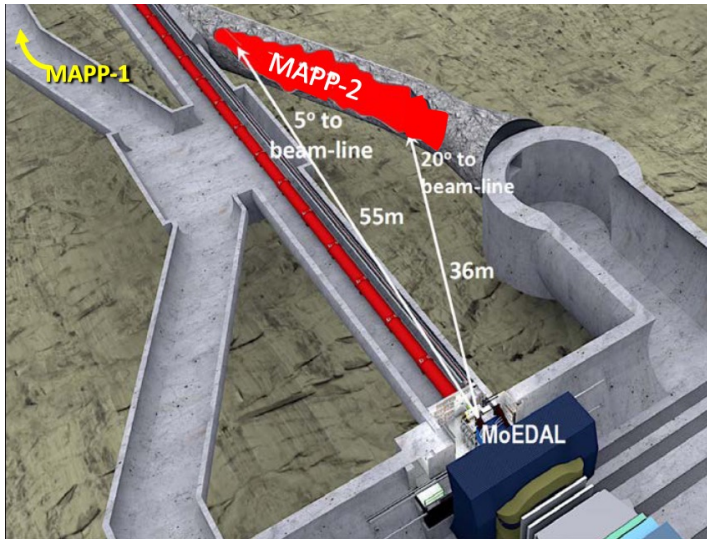
TIM (@VME-3) showing fiber input from the LHC



LEDs lighting up (ER, BR) after fiber input from LHC/ATLAS

DAQ



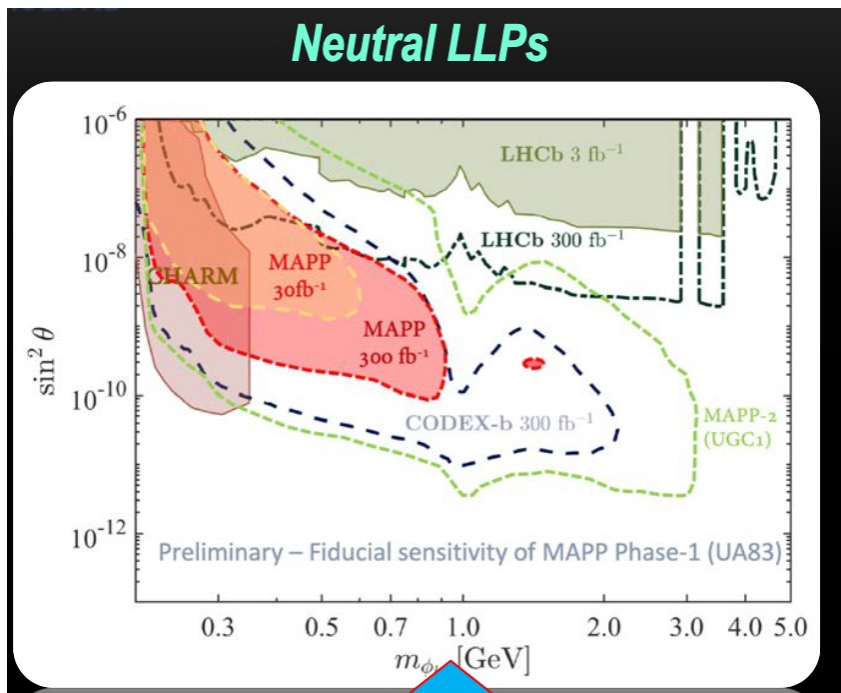


MAPP-2

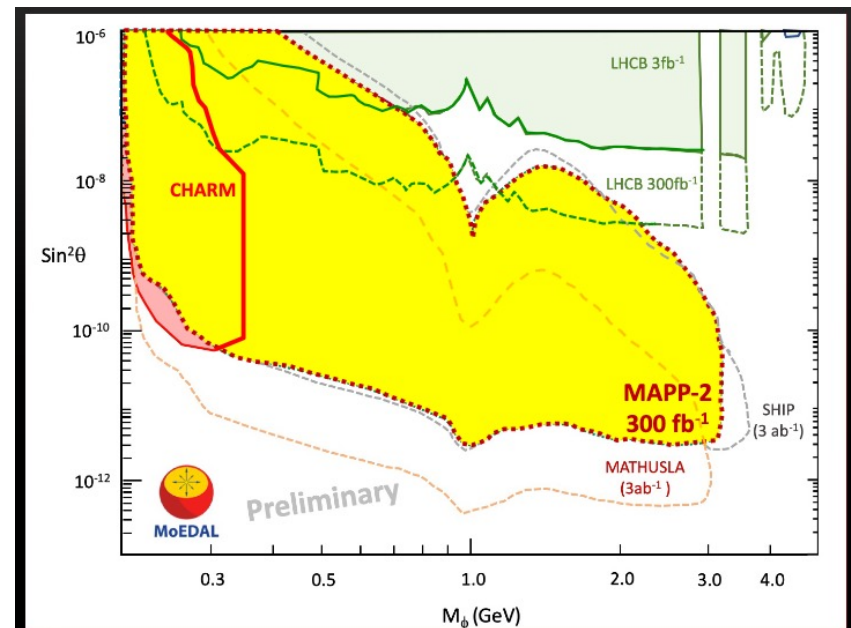


FASER

Benchmark: $B \rightarrow X_s \phi$



MAPP-1 LLP Sensitivity

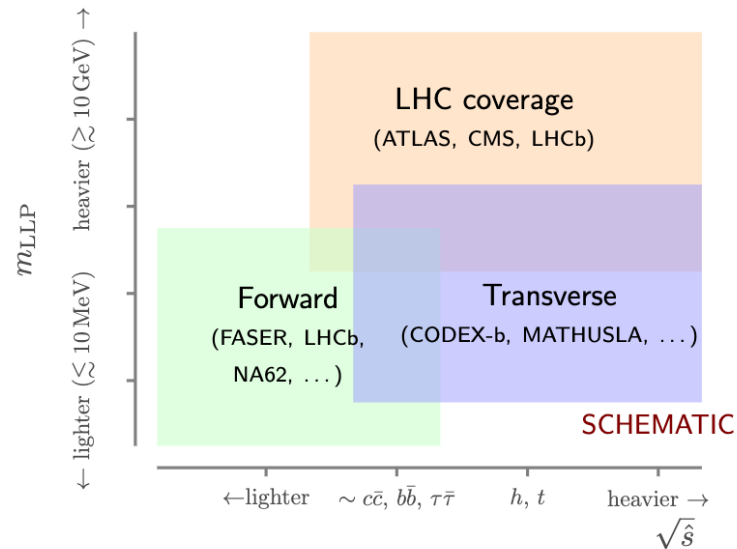


See Phys. Rev. D97 (1) (2018) 15023 for CODEX-b results.

MAPP-2 LLP Sensitivity

Dedicated detectors

Complementarity of different experiments searching for LLPs



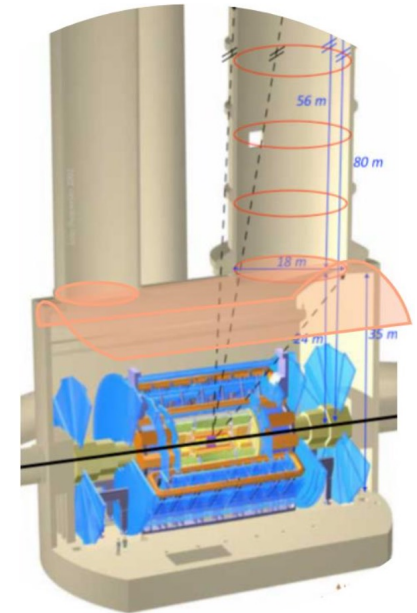
Regimes for which existing and proposed experiments have the most effective coverage (roughly):

1. ATLAS & CMS: Heavy LLPs ($m_{\text{LLP}} \gtrsim 10 \text{ GeV}$) for all lifetimes ($c\tau \lesssim 10^7 \text{ m}$).
2. LHCb: Short to medium lifetimes ($c\tau \lesssim 1 \text{ m}$) for light LLPs ($0.1 \text{ GeV} \lesssim m_{\text{LLP}} \lesssim 10 \text{ GeV}$).
3. Forward/beam dump detectors (FASER, NA62, SHiP): Medium to long lifetime regime ($0.1 \lesssim c\tau \lesssim 10^7 \text{ m}$) for light LLPs ($m_{\text{LLP}} \lesssim \text{few GeV}$), for low $\sqrt{\hat{s}}$ production channels.
4. Shielded, transversely displaced detectors (MATHUSLA, CODEX-b, AL3X): Relatively light LLPs² ($m_{\text{LLP}} \lesssim 10\text{--}100 \text{ GeV}$) in the long lifetime regime ($1 \lesssim c\tau \lesssim 10^7 \text{ m}$), and high $\sqrt{\hat{s}}$ production channels.

Dedicated detectors: ANUBIS

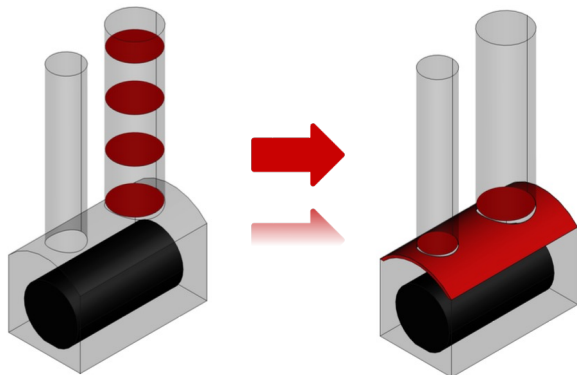
ANUBIS – AN Underground Belayed In-Shaft search experiment

- Proposal to instrument the PX14 shaft of ATLAS of the LHC at Point 1 (and/or possibly at Point 5)
 - > Would have 4 tracking stations equally spaced (between 20 and 80 m) in the shaft (and total area of detectors $2.3 \times 10^3 \text{ m}^2$)
 - > Each tracking station made of two layers of triplets separated by 1m
 - > Data taking during entire HL-LHC
 - > Close enough to ATLAS to participate in L1 trigger decision
- Practical concerns about the ability to quickly remove all 4 stations
- Alternative idea: instrument the ceiling of the ATLAS cavern
 - > Include stations in the two service shafts (PX14, PX16)
 - > Ceiling approximately 20m away from the ATLAS IP
 - > Cavern ceiling proposal shown to be more sensitive
 - > Larger active volume ($4.3 \times 10^4 \text{ m}^3$ vs $1.3 \times 10^4 \text{ m}^3$) and large detector area $\sim 10^3 \text{ m}^2$

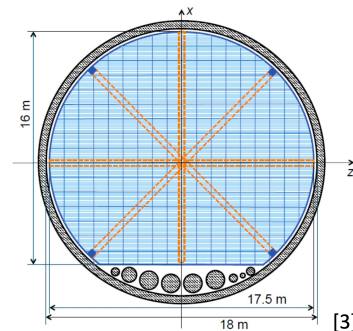


arXiv:1909.13022

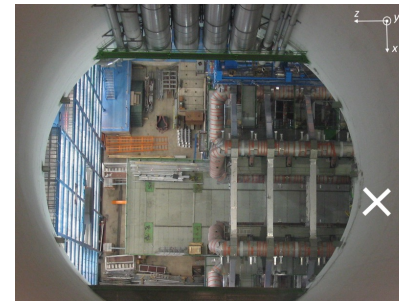
<https://twiki.cern.ch/twiki/bin/view/ANUBIS>



Original Vs Accepted



PX14 Shaft: Cross sectional view



PX14 Shaft



PX14 Shaft + Ceiling